

Find Real, Imaginary parts:

$$\frac{\nabla \epsilon}{\epsilon} = \frac{\nabla \epsilon_1 + i \nabla \epsilon_2}{\epsilon_1 + i \epsilon_2} = \frac{(\nabla \epsilon_1 + i \nabla \epsilon_2)(\epsilon_1 - i \epsilon_2)}{\epsilon_1^2 + \epsilon_2^2}$$

$$= \frac{\epsilon_1 \nabla \epsilon_1 + \epsilon_2 \nabla \epsilon_2}{\epsilon_1^2 + \epsilon_2^2} + i \frac{\epsilon_1 \nabla \epsilon_2 - \epsilon_2 \nabla \epsilon_1}{\epsilon_1^2 + \epsilon_2^2}$$

REAL PART:

$$[(\nabla S)^2 - \epsilon_1 \mu c^2] \bar{E}_0 =$$

$$= -\frac{1}{k_0} \left( \bar{E}_0 \cdot \frac{\epsilon_1 \nabla \epsilon_2 - \epsilon_2 \nabla \epsilon_1}{\epsilon_1^2 + \epsilon_2^2} \right) \nabla S + \frac{1}{k_0^2} \left[ \nabla^2 \bar{E}_0 + \nabla \left( \bar{E}_0 \cdot \frac{\epsilon_1 \nabla \epsilon_1 + \epsilon_2 \nabla \epsilon_2}{\epsilon_1^2 + \epsilon_2^2} \right) \right]$$

IMAGINARY PART:

$$- \epsilon_2 \mu c^2 \bar{E}_0 = \frac{2}{k_0} [(\nabla S \cdot \nabla) + \frac{1}{2} \nabla^2 S] \bar{E}_0$$

$$+ \frac{1}{k_0} \left( \bar{E}_0 \cdot \frac{\epsilon_1 \nabla \epsilon_1 + \epsilon_2 \nabla \epsilon_2}{\epsilon_1^2 + \epsilon_2^2} \right) \nabla S$$

$$+ \frac{1}{k_0^2} \nabla \left( \bar{E}_0 \cdot \frac{\epsilon_1 \nabla \epsilon_2 - \epsilon_2 \nabla \epsilon_1}{\epsilon_1^2 + \epsilon_2^2} \right)$$

EXACT TILL NOW (assumed:  $\mu(\vec{r}) = \mu$  real;  
 $S, \bar{E}_0$  real).

$\lambda \rightarrow 0$  ( $k_0 \rightarrow \infty$ ): Use only lowest order terms:

48  
cont.

REAL PART:

$$[(\nabla S)^2 - \epsilon_1 \mu c^2] \bar{E}_0 = 0 \quad (*)$$

↑ lowest order : 0th order

$$(\nabla S)^2 = \epsilon_1 \mu c^2 = \frac{\epsilon_1 \mu}{\epsilon_0 \mu_0} \equiv \tilde{n}^2$$

S is known as EIKONAL

and (\*) as EIKONAL APPROX.

BASIC EQ. OF GEOMETRIC OPTICS

S(C) = CONST defines GEOMETRICAL WAVE FRONT

( $\epsilon$  can be complex :)

Geometrical light rays = orthogonal trajectories to the geometrical wavefronts  $S(\vec{r}) = \text{constants}$ .

Oriented curves in direction of  $\vec{s} = \frac{\nabla S}{|\nabla S|}$

Eqs (3) and (4) to lowest order:

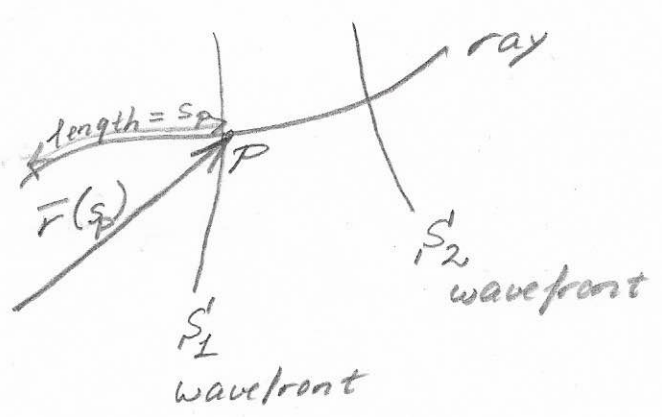
$$\vec{E}_0 \cdot \nabla S = 0 \quad \vec{H}_0 \cdot \nabla S = 0$$

So :

$\vec{E}_0$  and  $\vec{H}_0$  are orthogonal to the ray at every point.

Let  $\vec{r}(s)$  position vector of point P on ray

as a function of length of arcs  $s$  along the ray.



Then:

$$\vec{s} = \frac{d\vec{r}}{ds}$$

$$\frac{\nabla S}{|\nabla S|} = \vec{s} = \frac{d\vec{r}}{ds} \quad \text{and} \quad (\nabla S)^2 = \epsilon_1 \mu c^2 = \frac{\epsilon_1 \mu}{\epsilon_0 \mu_0}$$

note:  
 $\vec{n} = n\vec{r}$   
for  $\epsilon$  real

$$\text{So : } \nabla S = \frac{d\vec{r}}{ds} \cdot |\nabla S| = \frac{d\vec{r}}{ds} \cdot \frac{\sqrt{\epsilon_1 \mu}}{\sqrt{\epsilon_0 \mu_0}} = \frac{d\vec{r}}{ds} \cdot \vec{n}$$

$\frac{d\vec{r}}{ds} \vec{n} = \nabla S$

: Equation of Ray

Two neighboring wavefronts:

$$S'(r) = S_1 \quad S(r) = S_1 + ds$$

$$\frac{dS}{ds} = \nabla S \cdot \frac{d\vec{r}}{ds} = \tilde{n} \cdot \vec{s} = \tilde{n} \quad (*)$$

$\frac{dS}{ds}$   $\uparrow$   
 distance along ray between wavefronts  
 = distance along wavefront normal

DEFINE:  $\int_{\text{Ray II}} \tilde{n} ds = \text{optical length along ray}$   
 $\int_{\text{Ray}} n_r ds$  when  $\epsilon$  real

Optical length between points  $P_1, P_2$

on ray =  $\int_{P_1}^{P_2} \tilde{n} ds = \int_{P_1}^{P_2} dS' = S'(P_2) - S'(P_1)$   
 (\*)

For  $\epsilon$  real:  $c/\tilde{n} = c/n_r = v = \text{phase velocity of light}$

OPTICAL LENGTH EQUAL TO TIME FOR LIGHT TO TRAVEL  $P_1 \rightarrow P_2$  TIMES  $c \leftarrow$  vacuum speed

$$\boxed{S'(P_2) - S'(P_1) = \int_{P_1}^{P_2} \tilde{n} ds = \int_{P_1}^{P_2} c dt}$$

time needed for light to travel distance  $ds$  along ray

$$\frac{d\vec{r}}{ds} = \frac{\nabla S}{\tilde{n}}$$

$\uparrow$  tangent of curve (ray)  $\vec{r}(s)$ .  
 $\uparrow$  unit vector perpendicular to geometric wave front

We could solve for  $S$  and then solve this 1st order eq. for  $\vec{r}(s)$ .  
 Instead: derive 2nd order eq. for  $\vec{r}(s)$ :

$$\tilde{n} \frac{d\vec{r}}{ds} = \nabla S$$

$$\rightarrow \frac{d}{ds} \left( \tilde{n} \frac{d\vec{r}}{ds} \right) = \frac{d}{ds} (\nabla S) = \left( \frac{d\vec{r}}{ds} \cdot \nabla \right) \nabla S$$

you check!  $\rightarrow$

$$\begin{aligned} & \parallel \left( \frac{1}{\tilde{n}} \nabla S \cdot \nabla \right) \nabla S \\ & \parallel \frac{1}{2\tilde{n}} \nabla (\nabla S \cdot \nabla S) \\ & \parallel \frac{1}{2\tilde{n}} \nabla (\tilde{n}^2) \\ & \parallel \nabla \tilde{n} \end{aligned}$$

$$\frac{d}{ds} \left( \tilde{n} \frac{d\vec{r}}{ds} \right) = \nabla \tilde{n}$$

IN HOMOGENEOUS MEDIUM:

$$\frac{d}{ds} \left( \tilde{n} \frac{d\vec{r}}{ds} \right) = 0$$

$$\parallel$$

$$\tilde{n} \frac{d^2 \vec{r}}{ds^2} \Rightarrow \frac{d^2 \vec{r}}{ds^2} = 0 \quad ; \quad \vec{r} = s\vec{a} + \vec{b}$$

Light rays are straight lines

Now: Equation of motion for ray:

$$\frac{d}{ds} \left( \tilde{n} \frac{d\vec{r}}{ds} \right) = \nabla \tilde{n}$$

$$d\underline{t} \equiv ds \frac{\tilde{n}}{c}$$

$$\frac{\tilde{n}}{c} \frac{d}{d\underline{t}} \left( \tilde{n} \frac{\tilde{n}}{c} \frac{d\vec{r}}{d\underline{t}} \right) = \nabla \tilde{n}$$

$$\frac{d}{d\underline{t}} \left( \tilde{n}^2 \frac{d\vec{r}}{d\underline{t}} \right) = c^2 \frac{\nabla \tilde{n}}{\tilde{n}}$$

$$\left( \nabla(\tilde{n}^2) \cdot \frac{d\vec{r}}{d\underline{t}} \right) \frac{d\vec{r}}{d\underline{t}} + \tilde{n}^2 \frac{d^2 \vec{r}}{d\underline{t}^2} = \frac{c^2}{2} \frac{\nabla(\tilde{n}^2)}{\tilde{n}^2}$$

$$\left( \frac{\nabla(\tilde{n}^2)}{\tilde{n}^2} \cdot \frac{d\vec{r}}{d\underline{t}} \right) \frac{d\vec{r}}{d\underline{t}} + \frac{d^2 \vec{r}}{d\underline{t}^2} = \frac{c^2}{2} \frac{\nabla(\tilde{n}^2)}{\tilde{n}^4} \quad (*)$$

Remember:  $\tilde{n}^2 = \frac{\epsilon_1 \mu}{\epsilon_0 \mu_0}$

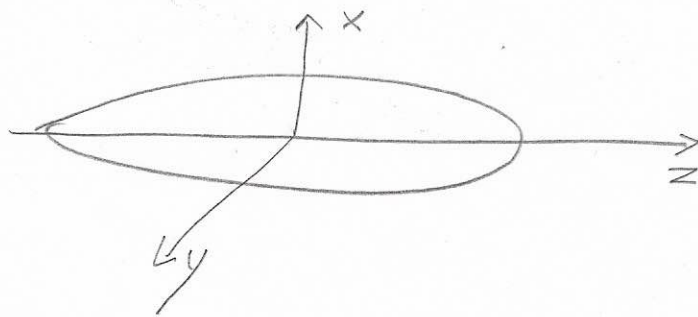
$$\epsilon = \epsilon_0 (1 + \chi) = \epsilon_0 + N \alpha \epsilon_0$$

atomic density

$$\epsilon_0 \cdot \text{atomic polarizability} = \frac{e^2}{m} \frac{1}{\omega_0^2 - \omega^2 - i\omega\gamma}$$

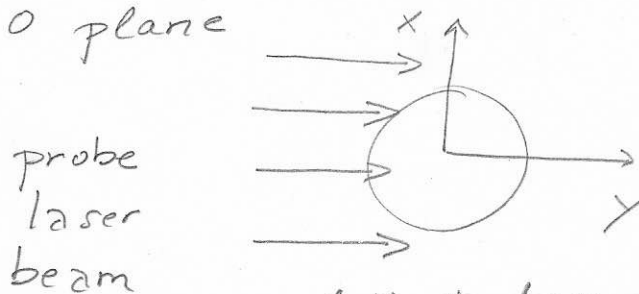
Set  $\mu = \mu_0$ :

Let  $z=0$ :  $\tilde{n}^2 = 1 + \text{Re}(N\alpha) = 1 + A e^{-(x^2+y^2)/(2\delta^2)}$



For: Non-condensed atom cloud  
( $\delta^2 \propto k_B T$ )  
or small condensate

In  $z=0$  plane



probe laser beam

What happens?

( $\delta = \sqrt{\frac{\hbar}{M\omega}}$  = size of ground state in harmonic oscillator)

$$\nabla(\tilde{n}^2) = -A \frac{2\vec{r}_\perp}{2\delta^2} e^{-r_\perp^2/(2\delta^2)} \quad \text{where}$$

$$\vec{r}_\perp = (x, y)$$

INTO (\*) :

$$\left( \frac{-A \bar{r}_\perp}{b^2} \frac{e^{-r}}{1+Ae^{-r}} \cdot \frac{d\bar{r}_\perp}{dt} \right) \frac{d\bar{r}_\perp}{dt} + \frac{d^2 \bar{r}_\perp}{dt^2}$$

$$= \frac{c^2}{2} \left( \frac{-A \bar{r}_\perp}{b^2} \frac{e^{-r}}{(1+Ae^{-r})^2} \right) ;$$

$$0 = \frac{d^2 \bar{r}_\perp}{dt^2} - \frac{A}{2b^2} \cdot \frac{d(r_\perp^2)}{dt} \frac{e^{-r}}{1+Ae^{-r}} \frac{d\bar{r}_\perp}{dt} + \frac{c^2 A}{2b^2} \frac{e^{-r}}{(1+Ae^{-r})^2} \bar{r}_\perp$$

Let  $\tau = \omega t$  unitless :

$$0 = \frac{d^2 \bar{r}_\perp}{d\tau^2} - \frac{A}{2b^2} \frac{d(r_\perp^2)}{d\tau} \frac{e^{-r_\perp^2/(2b^2)}}{1+Ae^{-r_\perp^2/(2b^2)}} \frac{d\bar{r}_\perp}{d\tau}$$

$$+ \frac{c^2}{\omega^2} \frac{A}{2b^2} \frac{e^{-r_\perp^2/(2b^2)}}{(1+Ae^{-r_\perp^2/(2b^2)})^2} \bar{r}_\perp$$

$$= \frac{1}{k_0} = \left( \frac{\lambda_0}{2\pi} \right)^2$$

Put back in  $\tilde{n}^2$ :

$$0 = \ddot{\vec{r}}_{\perp}(\tau) - \left[ \frac{\tilde{n}^2 - 1}{\tilde{n}^2} \frac{1}{6^2} \vec{r}_{\perp}(\tau) \cdot \dot{\vec{r}}_{\perp}(\tau) \right] \dot{\vec{r}}_{\perp}(\tau) + \frac{\lambda_0^2}{8\pi^2} \frac{1}{6^2} \frac{\tilde{n}^2 - 1}{\tilde{n}^4} \vec{r}_{\perp}(\tau)$$

Equation valid only for Gaussian distrib.

Gaussian N



Parabolic distribution:

$$N = A - B(x^2 + y^2 + \epsilon z^2)$$

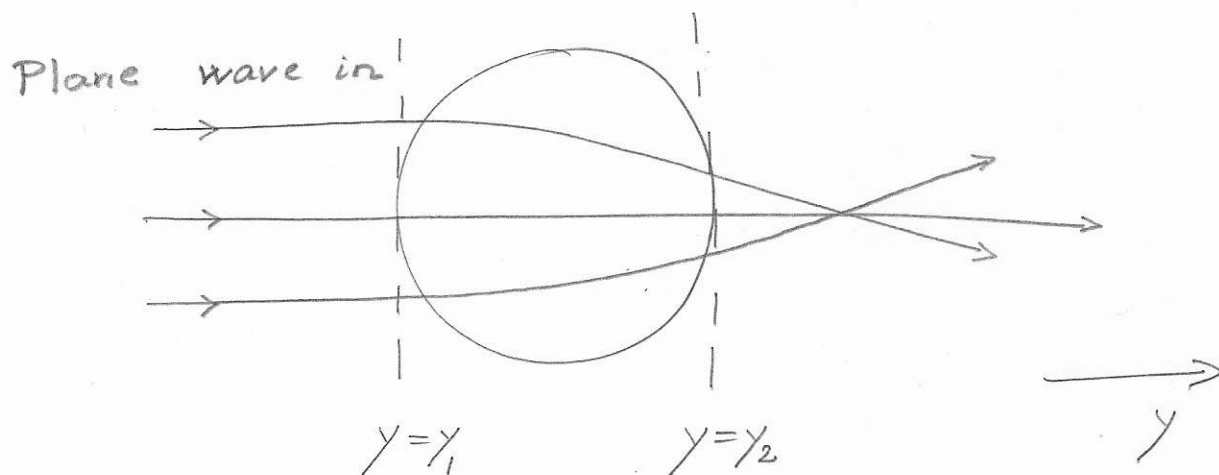
for large condensate

Solve for  $\vec{r}_{\perp}(\tau)$ : then we have ray trajectory

$\vec{r}_{\perp}(\tau_1) = P_1$        $\vec{r}_{\perp}(\tau_2) = P_2$

$$S(P_2) - S(P_1) = \int_{P_1}^{P_2} \tilde{n} ds = \int_{P_1}^{P_2} c d\underline{t}$$

$$\underbrace{e^{i k_0 S(\vec{r})}}_{\text{phase}} \underbrace{e^{-i \omega t}}_{\text{time}} = \frac{c}{\omega} \int_{P_1}^{P_2} d\tau = \frac{\lambda_0}{2\pi} \int_{P_1}^{P_2} d\tau$$



In plane  $y=y_1$ :  $\vec{r}_\perp(\tau) = \begin{pmatrix} x \\ y_1 \end{pmatrix}; \tau=0$

Cloud is hit by plane wave:

$$\vec{E}_0(\vec{r}) e^{ik_0 S(\vec{r})} e^{-i\omega t} = E_2 e^{ik_0 y_1} e^{-i\omega t}$$

$\varphi$   
wavefront = plane  
 $y=y_1$

$$s_0: \vec{s} = \frac{d\vec{r}}{ds} = \hat{e}_y$$

$$\frac{d\vec{r}}{cdt} = \frac{\omega}{c} \frac{d\vec{r}}{d\tau} = \frac{2\pi}{\lambda_0} \dot{\vec{r}}(\tau)$$

Initial condition }  $\vec{r}_\perp = \begin{pmatrix} x_1 \\ y=y_1 \end{pmatrix}$   $\dot{\vec{r}}_\perp = \frac{\lambda_0}{2\pi} \begin{pmatrix} 0 \\ 1 \end{pmatrix}$   
in plane  
 $y=y_1$  at  $\tau=0$

Propagate each ray (defined by  $x_1$ ) to plane  
 $y=y_2$

Propagate to plane  $y = y_2$  : gives phase

Note  $x_2$  and  $\tau_2$  values for each ray.

At  $(x_2, y_2)$  we have phase factor  $e^{ik_0 \Delta S} e^{ik_0 x_2}$

$$\Delta S = S_2 - S_1 = \int_{P_1}^{P_2} \tilde{n} ds = \frac{\lambda_0}{2\pi} \int_{P_1}^{P_2} d\tau$$

$$= \frac{\lambda_0}{2\pi} \tau_2 = \frac{1}{k_0} \tau_2$$

at  $y = y_1$

This defines phase distribution

in plane  $y = y_2$

and thereby wavefront distortion caused by atom cloud.

From here we propagate through imaging system using Fourier optics

How about  $\vec{E}_0(\vec{r})$  : absorption

# ABSORPTION:

IMAGINARY PART in nearly homogeneous medium:

$$-\epsilon_2 \mu c^2 \bar{E}_0 = \frac{2}{k_0} (\nabla S \cdot \nabla) \bar{E}_0$$

$$= \frac{2}{k_0} \left( \tilde{n} \frac{d\bar{r}}{ds} \cdot \nabla \right) \bar{E}_0$$

Again:  $d\underline{t} = ds \cdot \frac{\tilde{n}}{c}$

$$-\epsilon_2 \mu c^2 \bar{E}_0 = \frac{2}{k_0} \frac{\tilde{n}^2}{c} \left( \frac{d\bar{r}}{d\underline{t}} \cdot \nabla \right) \bar{E}_0$$

$$-\epsilon_2 \mu c^2 \bar{E}_0 = \frac{2}{k_0} \frac{\tilde{n}^2}{c} \frac{d\bar{E}_0}{d\underline{t}}$$

$$\Rightarrow \frac{d\bar{E}_0}{d\underline{t}} = - \frac{\epsilon_2 \mu c^2}{2} \cdot \frac{k_0 c}{\tilde{n}^2} \bar{E}_0$$

$$\Rightarrow \frac{d\bar{E}_0}{d\underline{\tau}} = - \frac{\epsilon_2 \mu c^2}{2 \tilde{n}^2} \left( \frac{k_0 c}{\omega} \right) \bar{E}_0 = - \frac{\epsilon_2 \mu}{\epsilon_0 \mu_0} \frac{1}{2 \tilde{n}^2} \bar{E}_0$$

$\uparrow$   
 $\underline{\tau} = \omega t$

$\uparrow$   
 $k_0 / k_0 = 1$

$$\bar{E}_0(\underline{\tau}) = \bar{E}_0(\underline{\tau}=0) e^{-\gamma \underline{\tau}} ;$$

if  $\epsilon_2 / \epsilon_1$   
indep. of  $\bar{r}$   
ex.: HOMOGENEOUS  
MEDIUM

$$\gamma = \frac{\epsilon_2 \mu}{\epsilon_0 \mu_0} \frac{1}{2 \tilde{n}^2}$$

$$= \frac{\epsilon_2 \mu}{\epsilon_0 \mu_0} \frac{1}{2} \frac{\epsilon_0 \mu_0}{\epsilon_1 \mu}$$

$$= \frac{1}{2} \frac{\epsilon_2}{\epsilon_1}$$

If  $\epsilon_2/\epsilon_1$  is spatially dependent  
we have for linearly polarized light  
along  $z$ :

$$\frac{dE_{0z}}{d\tau} = -\frac{1}{2} \frac{\epsilon_2}{\epsilon_1} E_{0z} \Rightarrow \frac{dE_{0z}}{E_{0z}} = -\frac{1}{2} \frac{\epsilon_2}{\epsilon_1} d\tau.$$

Then

$$\int_{P_1}^{P_2} \frac{dE_{0z}}{E_{0z}} = -\frac{1}{2} \int_{\tau_1}^{\tau_2} \frac{\epsilon_2(\vec{r}(\tau))}{\epsilon_1(\vec{r}(\tau))} d\tau, \text{ where}$$

$$\begin{aligned} \int_{P_1}^{P_2} \frac{dE_{0z}}{E_{0z}} &= \ln E_{0z}(P_2) - \ln E_{0z}(P_1) \\ &= \ln \frac{E_{0z}(P_2)}{E_{0z}(P_1)}, \text{ resulting in} \end{aligned}$$

$$E_{0z}(P_2) = e^{-\frac{1}{2} \int_{\tau_1}^{\tau_2} d\tau \epsilon_2(\vec{r}(\tau)) / \epsilon_1(\vec{r}(\tau))} E_{0z}(P_1)$$

Include this in wavefront at plane  $y = y_2$ :

$$E_{0z}(P_2) e^{ik_0 S(\tau_2)} e^{-i\omega t}$$

When is EIKONAL APPROX  
RAY OPTICS

VALID ? :

$$\lambda \rightarrow 0 \quad \left( \frac{1}{k_0} \rightarrow 0 \right) ;$$

Need dimensionless quantity :

$$\frac{1}{k_0} \frac{\nabla \epsilon}{\epsilon} \ll 1$$

||

$$\frac{\lambda_0}{2\pi} \frac{\nabla \epsilon}{\epsilon} \Rightarrow \frac{\nabla \epsilon}{\epsilon} \ll \frac{1}{\lambda_0}$$

$$\frac{1}{k_0} \frac{\nabla \epsilon_0}{\epsilon_0} \ll 1 \Rightarrow \frac{\nabla \epsilon_0}{\epsilon_0} \ll \frac{1}{\lambda_0}$$



cannot treat focal points