Physics 253a Problem set 3

Due Thursday October 20, 2022

Problem 1. Consider in *D*-dimensional spacetime a free scalar field theory with the Lagrangian density

$$\mathcal{L} = -\frac{1}{2}\eta^{\mu\nu}\partial_{\mu}\phi\partial_{\nu}\phi - \frac{1}{2}m^{2}\phi^{2}.$$
 (1)

In class we have derived the Fourier representation of the time-ordered 2-point function

$$G(x,y) \equiv \langle \Omega | \mathbf{T}\hat{\phi}(x)\hat{\phi}(y) | \Omega \rangle = -i \int \frac{d^D k}{(2\pi)^D} \frac{e^{ik \cdot (x-y)}}{k^2 + m^2 - i\epsilon}$$
(2)

where $k \cdot (x - y) = \vec{k} \cdot (\vec{x} - \vec{y}) - k^0 (x^0 - y^0)$ and $k^2 = -(k^0)^2 + \vec{k}^2$, as well as the Euclidean 2-point function

$$G_E(x_E, y_E) \equiv \langle \hat{\phi}_E(x_E) \hat{\phi}_E(y_E) \rangle = \int \frac{d^D k_E}{(2\pi)^D} \frac{e^{ik_E \cdot (x_E - y_E)}}{k_E^2 + m^2},$$
(3)

where $k_E \cdot (x_E - y_E) = \sum_{\mu=1}^{D} k_E^{\mu} (x_E - y_E)^{\mu}$ and $k_E^2 = \sum_{\mu=1}^{D} (k_E^{\mu})^2$, and $\hat{\phi}_E(x_E)$ is related to $\hat{\phi}(x)$ by the analytic continuation in x^0 that takes $x^0 \to -ix_E^D$.

For general D, the k-integral can be expressed in terms of Bessel functions. When D is an odd integer, the result is an elementary function. In this problem, you are asked to consider the D = 3 case, and write an elementary expression for G(x, y), $G_E(x_E, y_E)$, as well as the Wightman function

$$W(x,y) \equiv \langle \Omega | \hat{\phi}(x) \hat{\phi}(y) | \Omega \rangle.$$
(4)

Explain how W(x, y) is related to $G_E(x_E, y_E)$ by analytic continuation, and how W(x, y) differs from W(y, x) for time-like separated x, y.

To be sure that you understand what's going on, make a plot for both the real and imaginary part of W(x, 0) as a function of (real time) x^0 in the case m = 0 for some fixed spatial coordinate $\vec{x} \neq 0$. **Problem 2.** Consider in *D*-dimensional spacetime a free *complex* scalar field theory with the Lagrangian density

$$\mathcal{L} = -\eta^{\mu\nu}\partial_{\mu}\varphi^*\partial_{\nu}\varphi - m^2\varphi^*\varphi, \qquad (5)$$

where $\varphi = \phi_1 + i\phi_2$, $\varphi^* = \phi_1 - i\phi_2$. The U(1) symmetry that rotates the phase of φ is associated with a Noether current operator, of the form¹

$$\hat{j}_{\mu}(x) = -i : \hat{\varphi}^{*}(x) \overleftrightarrow{\partial}_{\mu} \hat{\varphi}(x) :$$
(7)

where $f \overleftrightarrow{\partial}_{\mu} g \equiv f \partial_{\mu} g - (\partial_{\mu} f) g$.

(a) Verify that \hat{j}^{μ} is conserved, namely $\partial_{\mu}\hat{j}^{\mu} = 0$ as an operator.

(b) Calculate the time-ordered 3-point correlation function

$$H^{\mu}(x_{1}, x_{2}, x_{3}) \equiv \langle \Omega | \mathbf{T} \hat{\varphi}^{*}(x_{1}) \varphi(x_{2}) \hat{j}^{\mu}(x_{3}) | \Omega \rangle$$

=
$$\int \frac{d^{D} k_{1}}{(2\pi)^{D}} \frac{d^{D} k_{2}}{(2\pi)^{D}} \widetilde{H}^{\mu}(k_{1}, k_{2}) e^{ik_{1} \cdot x_{1} + ik_{2} \cdot x_{2} - i(k_{1} + k_{2}) \cdot x_{3}}.$$
 (8)

(c) Calculate

$$\frac{\partial}{\partial x_3^{\mu}} H^{\mu}(x_1, x_2, x_3) = \int \frac{d^D k_1}{(2\pi)^D} \frac{d^D k_2}{(2\pi)^D} (-i(k_1 + k_2)_{\mu}) \widetilde{H}^{\mu}(k_1, k_2) e^{ik_1 \cdot x_1 + ik_2 \cdot x_2 - i(k_1 + k_2) \cdot x_3}.$$
 (9)

Does the result vanish? Explain whether/how it is compatible with your answer to (a).

$$\hat{j}_{\mu}(x) = -i \lim_{y \to x} \left[\hat{\varphi}^{*}(y) \overleftrightarrow{\partial}_{\mu} \hat{\varphi}(x) - \left\langle \hat{\varphi}^{*}(y) \overleftrightarrow{\partial}_{\mu} \hat{\varphi}(x) \right\rangle \right].$$
(6)

¹An equivalent expression for \hat{j}_{μ} in terms of $\hat{\phi}$ is

Problem 3. In class we introduced the 1-particle state $|\vec{k}\rangle$ for a massive particle with no internal degree of freedom, normalized according to $\langle \vec{k} | \vec{k}' \rangle = \delta^{D-1}(\vec{k} - \vec{k}')$, on which the Lorentz transformation acts by

$$U(\Lambda)|\vec{k}\rangle = \sqrt{\frac{(\Lambda k)^0}{k^0}}|\overrightarrow{\Lambda k}\rangle,\tag{10}$$

where $k^0 = \omega_{\vec{k}} \equiv \sqrt{\vec{k}^2 + m^2}$. Recall that for infinitesimal Lorentz transformation $\Lambda^{\mu}{}_{\nu} = \delta^{\mu}{}_{\nu} + \omega^{\mu}{}_{\nu}, U(\Lambda) = 1 + \frac{i}{2}\omega_{\mu\nu}\hat{J}^{\mu\nu}$. Restricting to the case of D = 4 spacetime dimensions, the usual angular momentum operator $\vec{J} = (J_1, J_2, J_3)$ is related by $\hat{J}_{ij} = \sum_{k=1}^3 \epsilon_{ijk} J_k$.

Now consider a two-particle state $|\vec{k_1}, \vec{k_2}\rangle$. For now you may assume that the two particles are not interacting, and that $U(\Lambda)$ acts by the obvious generalization of (10)²

$$U(\Lambda)|\vec{k}_1, k_2\rangle = \sqrt{\frac{(\Lambda k_1)^0 (\Lambda k_2)^0}{k_1^0 k_2^0}} |\overrightarrow{\Lambda k_1}, \overrightarrow{\Lambda k_2}\rangle.$$
(11)

Assuming the two particles are identical bosons, it is natural to normalize the two-particle state according to

$$\langle \vec{k}_1, \vec{k}_2 | \vec{k}_1', \vec{k}_2' \rangle = \delta^3(\vec{k}_1 - \vec{k}_1')\delta^3(\vec{k}_2 - \vec{k}_2') + \delta^3(\vec{k}_1 - \vec{k}_2')\delta^3(\vec{k}_2 - \vec{k}_1').$$
(12)

We can consider an alternative basis $|E, \vec{P}, \ell, m\rangle$ for the two-particle state, that are eigenstates with respect to the total energy-momentum vector (of eigenvalue (E, \vec{P})), and with respect to the angular momentum squared $\vec{J}^2 = \ell(\ell+1)$ and the angular momentum component $J_3 = m$ defined in the center-of-mass frame. We will normalize these states according to the convention

$$\langle E, \vec{P}, \ell, m | E', \vec{P}', \ell', m' \rangle = \delta(E - E') \delta^3(\vec{P} - \vec{P}') \delta_{\ell\ell'} \delta_{mm'}.$$
(13)

It is easy to see that the overlap between the two kinds of basis states are of the form

$$\langle \vec{k}_1, \vec{k}_2 | E, \vec{P}, \ell, m \rangle = \delta(\omega_{\vec{k}_1} + \omega_{\vec{k}_2} - E) \delta^3(\vec{k}_1 + \vec{k}_2 - \vec{P}) f_{\ell m}(E, \vec{P}; \hat{k}_1)$$
(14)

where the function $f_{\ell m}$ depends only on E, \vec{P} , and the unit vector $\hat{k}_1 \equiv \vec{k}_1/|\vec{k}_1|$. Your task is to determine $f_{\ell m}$ in the special case $\vec{P} = 0$. In particular, show that

$$f_{\ell m}(E,0;\hat{k}_1) = \mathcal{N}(E)Y_{\ell m}(\hat{k}_1),$$
(15)

where $Y_{\ell m}$ are the spherical harmonics, and determine the function $\mathcal{N}(E)$.

 $^{^{2}}$ We will see in class soon that this is true for the so-called in- and out- asymptotic scattering states in an interacting quantum field theory as well.

Problem 4. A key fact concerning the Lorentz transformation property of the 1-particle state, as discussed in class, is that the action of $U(\Lambda)$ on the 1-particle state is determined through that of U(W), where W is a Lorentz transformation that leaves the momentum of the particle invariant. The set of all such W's for a given particle momentum is called the "little group". In this problem you will study the little group of a *massless* particle in D = 4 spacetime dimensions.

(a) We begin with a light-like reference momentum k_R^{μ} ,

$$(k_R^0, k_R^1, k_R^2, k_R^3) = (E, 0, 0, E).$$
(16)

For any spatial momentum 3-vector \vec{k} , and the corresponding light-like 4-momentum vector $(k^0 = |\vec{k}|, \vec{k})$, we can construct a "standard" Lorentz transformation $L(\vec{k})$ that takes k_R to k as

$$L(\vec{k}) = M(\vec{k}) \cdot R_1, \tag{17}$$

where

$$(M(\vec{k}))^{\mu}{}_{\nu} = \delta^{\mu}_{\nu} + \frac{(k - k_R)^{\mu}(k - k_R)_{\nu}}{k \cdot k_R},$$
(18)

and R_1 is the reflection in the x^1 -direction. Verify that (17) is indeed a Lorentz transformation of the desired property.

(b) $W = (L(\Lambda k_R))^{-1}\Lambda$, which leaves k_R invariant (i.e. $Wk_R = k_R$) for any generic Lorentz transformation Λ , is an element of the little group. Find an example of W that is not merely a rotation about the x^3 axis nor a reflection.

Remark: The structure of the massless little group you find here, as we will see in a few weeks, is closely related to why the photon is necessarily described by a gauge theory.